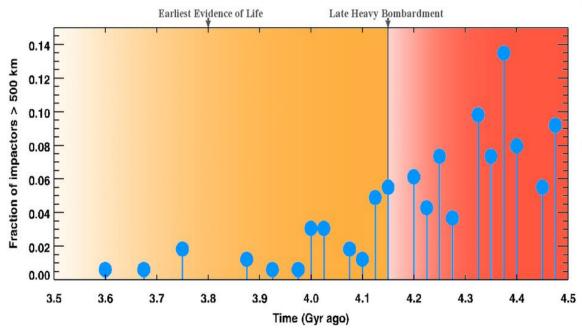


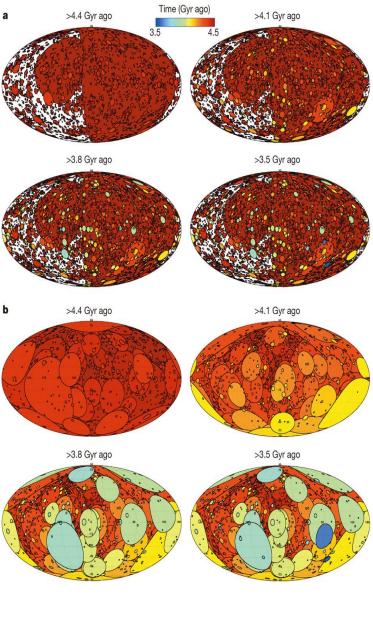
Jane Greaves

University of St Andrews, Scotland

# the hostile early Earth

 evidence of life only after heavy bombardment (e.g. Marchi et al. 2014)

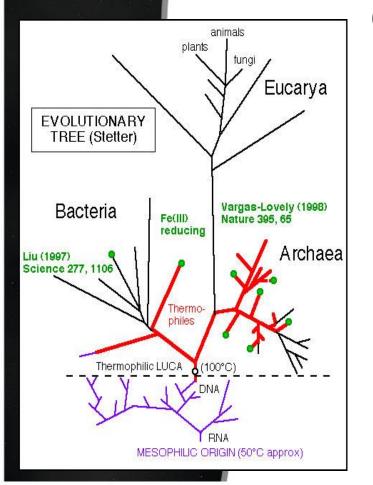




# impacts and energy for life

• thermal basins (stability? expansion?)

(Brochier & Philippe 2002; Flaspohler 2012)





# comet-belt histories

10-

 $10^{-2}$ 

 $10^{-3}$ 

10-4

 $10^{-5}$ 

10-6

10-7

10

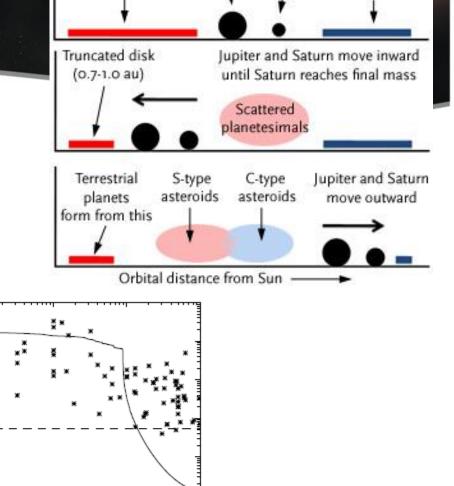
100

Time, Myr

is the solar system peculiar?

(Booth et al. 2009; Walsh et al. 2011)

1000



upiter

Saturn

Disk of water-

and carbon-rich

planetesimals

Disk of rocky

planetesimals

Figure 1. Excess ratio versus time for 10  $\mu$ m (left) and 70  $\mu$ m (right). The solid line represents the emission from our model. The asterisks are observed discs and the dashed line shows the approximate observational limit.

10000

10<sup>2</sup>

10-1

10-2

10

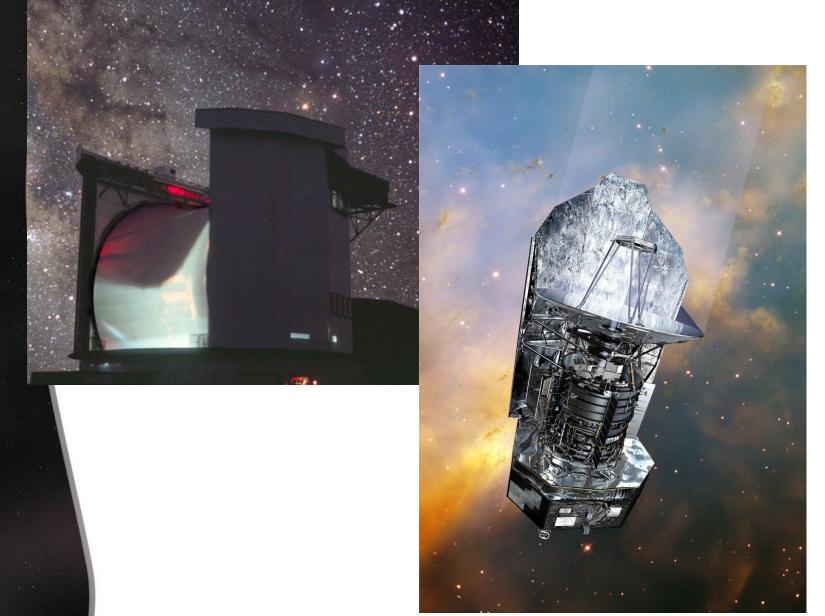
100

Time, Myr

1000

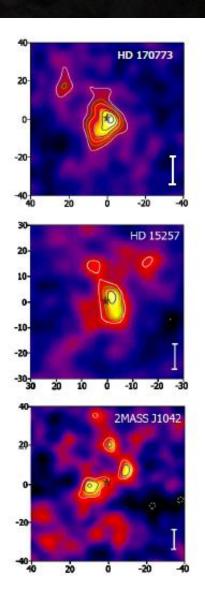
10000

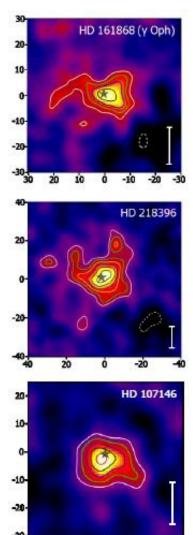
# observing exo-comet-belts

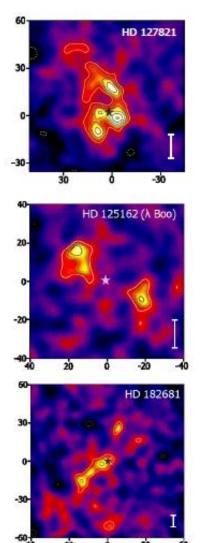


# Survey Of Nearby Stars

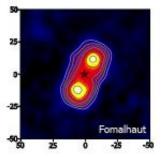


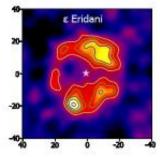








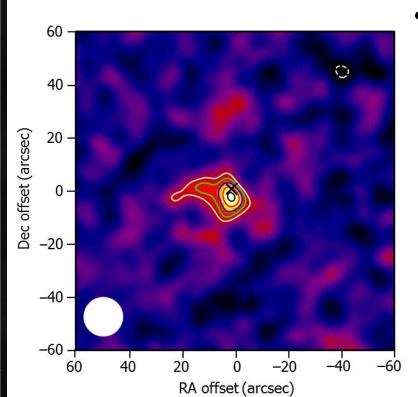


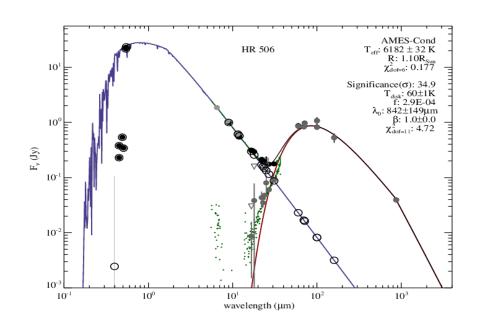


### comet belts on steroids...!

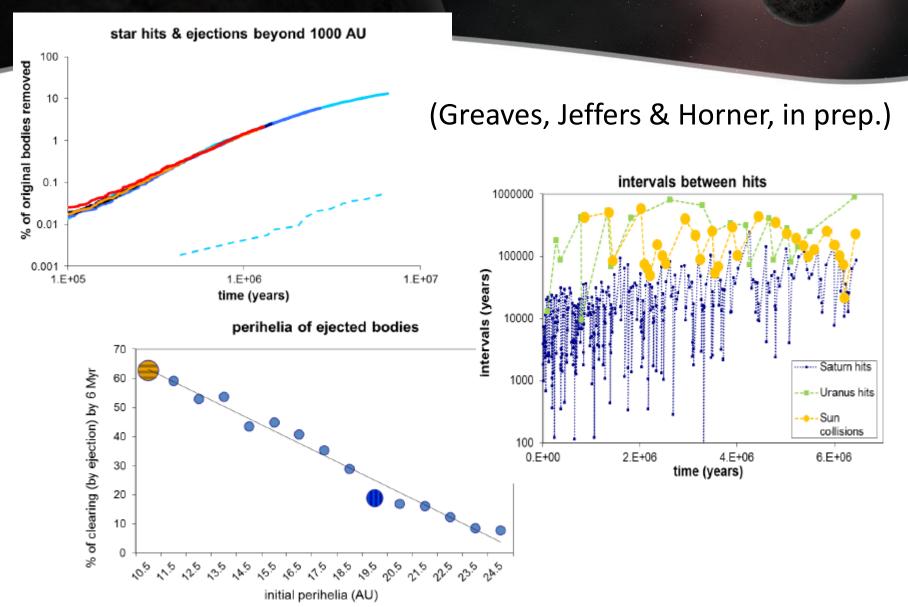
- q1 Eri (HD 10647) (→ Liseau et al. 2008)
  - dust belt of luminosity ~1000x solar
  - dust temperature ~1.5x vs. Kuiper Belt

• planet host star; age  $\approx 3.2 \pm 1.2$  Gyr (Bonfanti et al. 2015)



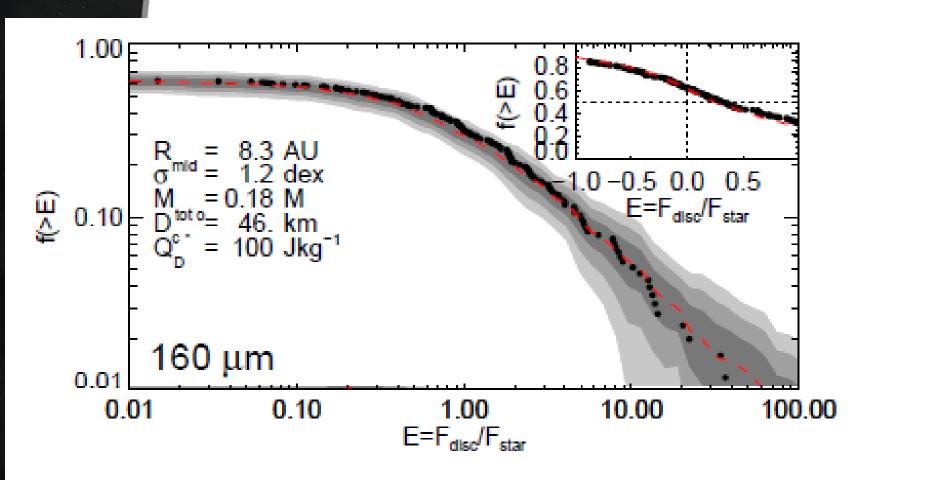


# impact simulations



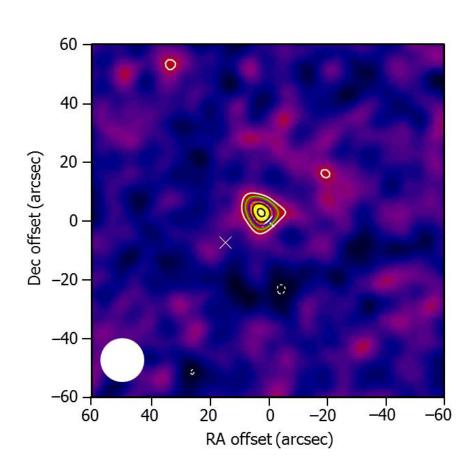
### debris around FGK stars

(Sibthorpe et al., in prep.)



# planets & comets at >> t<sub>Sun</sub>

- archetype is  $\tau$  Ceti
  - t ≈ 10 Gyr
  - 3+ low-massplanets
  - has a cometbelt withluminosity>> the Sun's



### what is needed...?

- compare catalogues of debris discs and planets
- simulate impacts for a variety of systems
- critique the outcomes:
  - ... lava worlds?
  - ...watery worlds?
  - ... Warm basins?