

SCIENTIFIC OPPORTUNITIES FOR A STARSHADE **WORKING WITH A** 2.4 M TELESCOPE AT L2

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Exoplanet Exploration Program

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NASA's recent Exoplanet Probe studies

- The discovery of exoEarths, via a space-based direct imaging mission, is a long-term priority for US astrophysics
- Two studies of "probe sized" exoplanet direct imaging missions: one coronagraph (Exo-C), one starshade (Exo-S)
 - Total mission cost targeted at \$1B (FY15 dollars)
 - New start in 2017
 - Compelling science must be beyond the expected ground capability at the time of mission
- Studies also intended as a design input to the exoplanet community to help formulate ideas for the next Decadal Survey in 2020

Exo-S Team Members

STDT

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Starshade Basics



- Contrast and IWA decoupled from telescope aperture size
- No outer working angle
- High throughput, broad wavelength bandpass
- High quality telescope not required
 - Wavefront correction unnecessary
- Retargeting requires long starshade slews (days to weeks)

Two Cost Constrained Exo-S Concepts

• Exo-S Dedicated Co-Launched Mission

- Starshade and telescope launch together to conserve cost
- Telescope: low-cost commercial Earth observer, 1.1 m diameter aperture
- Starshade: 30 m diameter
- Exo-S Rendezvous Mission
 - Starshade launches for a rendezvous with an existing telescope
 - Telescope: WFIRST/AFTA 2.4 m is adopted
 - Starshade: 34 m diameter
 - Orbit: Earth-Sun L2 (assumption for the purposes of the Exo-S study)
 - Three year Class C mission
 - Minimal impact to current mission design
 - No stringent requirements are imposed on the WFIRST/AFTA spacecraft
 - No new instrument, only modification to the existing coronagraph

WFIRST-AFTA + Starshade simulated image of Beta Canum Venaticorum plus solar system planets (8.44 pc, G0V)



Saturn

Hypothetical dust ring at 15 AU

Background galaxy

Image credit: M. Kuchner

Exo-S Science Goals



Simulated R=70 planet spectra for the 2.4-m mission, with three representative 10% error bars.

1.1-m mission cannot reach R=70 on small planets.

- Discover new exoplanets from giants down to Earth size
- 2. Characterize new planets with R=10 to 70 optical spectra
- Characterize known giant planets with R=70 spectra and constrain masses
- 4. Study planetary systems including circumstellar dust
 - Locate dust parent bodies
 - Evidence of unseen planets
 - Exozodi assessment for future missions

Key Capabilities

Wide-Field Imager, Integral Field Spectrograph, Guide Camera Already part of WFIRST coronagraph

425 – 602 nm	600 – 850 nm	706 – 1000 nm	
31%	43 %	51%	
Blue	G <mark>r</mark> een	Red	
IWA = 70 mas	IWA = 100 mas	IWA = 118 mas	
z = 50000 km	z = 35000 km	z = 30000 km	

- Contrast at inner working angle = 1×10^{-10}
- Total throughput = 28% (imager), 22% (IFS)
- Field-of-view = 10 arcsec (imager), 2 arcsec (IFS)

Observing Sequence

- 1. Schedule known giant planet observations
- 2. Fill gaps on sky with highest priority blind search target (Earths)
- 3. Repeat with lower priority targets (sub-Neptunes & Jupiters) until fuel or time limit (2 years) reached



55 stars visited. Uses \sim 9% of total WFIRST time.

12 known giant planets. Blind search targets: 28 Earths, 7 sub-Neptunes, 8 Jupiters

Expected Yield By Planet Type & Temperature



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Technical challenges

- Precise edge profile
 (~ 50 µm tolerance) required
 over large structure
- Knife-edge to prevent sunlight scattering into telescope
- On-orbit deployment of large structure



NASA / JPL / Princeton

- Precise alignment between starshade and telescope needed (± 1 meter tolerance). Soft requirement
- Full-scale end-to-end system test on the ground not possible

Contrast demonstrations

Caveat : lab & field demonstrations not completely accurate

Optical models with distortions monochromatic: 10⁻¹² Credit: S. Shaklan (JPL)





0.1% scale lab testing monochromatic: 10⁻¹⁰ Credit: J. Kasdin (Princeton)



 $\sim 1\%$ scale field testing September 23 HG Baseline 1.5 50% bandpass: 10⁻⁸ 4 x 10⁻⁸ -100Credit: Northrop Grumman "planet" Position [arcsec] 0 0.5 100 -300 -200 -100 100 200300 0 Position [arcsec]

Contrast [1E-7]

Precision petal manufacturing

Full-scale petal with edge profile for contrast < 10^{-10⁻}



Development of knife-edge to control edge scatter underway

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Deployment demonstration



Cost Estimates

- Cost and technology readiness analysis by Exo-S
 Team, JPL Team X, and Aerospace Corporation
- Setimated 2.4-m mission Phase A F cost: \$627M
 - Includes launch and modifications to WFIRST
- Aerospace Corp. raised no issues with technology schedule
 - To reach TRL-5 by 2017 for 5 key technology gaps

The cost information contained in this document is of a budgetary and planning nature and is intended for informational purposes only. It does not constitute a commitment on the part of JPL and Caltech.

Take-Away Message

WFIRST-AFTA can be leveraged for a unique and timely opportunity

- Mission with 34-m starshade can access exoEarths in the habitable zone for up to 50 unique target stars
- JPL now studying enhanced design with 40-meter starshade
- Minimal modification to WFIRST needed for starshade readiness
- Starshade technology is on track for new start by 2018, but not fully funded